

Principal Investigators:

Richard Aster aster@ees.nmt.edu

Philip Kyle kyle@nmt.edu

Both at New Mexico Institute of Mining and Technology, Socorro, NM, 87801

Experiment Name: Tomo Erebus

Seismic Tomography of Erebus Volcano, Antarctica

Daria Zandomenighi¹, Philip Kyle², Pnina Miller³, Catherine Snelson², Richard Aster²

¹ - International Centre for Theoretical Physics, Trieste, Italy; E-mail: dzandome@ictp.it

² - Department of Earth and Environmental Science, New Mexico Institute of Mining and Technology, Socorro

³ - IRIS PASSCAL Instrument Center, New Mexico Institute of Mining and Technology

Mount Erebus (77°32'S, 167°10'E; elevation 3794 meters) is the most active volcano in Antarctica and is well known for its persistent lava lake. The lake constitutes an “open window” into the conduit and underlying feeding system and offers a rare opportunity to observe a shallow convecting magmatic system.

Imaging and modeling of the internal structure of Erebus volcano are best done through compiling information from arrays of seismometers positioned strategically around the volcano. From these data, the three-dimensional (3-D) structure of the conduit can be pieced together. Building this 3-D model of Erebus was a main goal of the seismic tomographic experiment Tomo Erebus (TE). During the 2007–2008 austral field season, 23 intermediate-period seismometers were installed to contribute data, through the winter, for the passive-source aspect of the experiment. One year later, 100 three-component short-period stations were deployed to record 16 chemical blasts (see Figure 1).

These networks will help scientists update older 2-D models of the magmatic system underlying the lava lake [Dibble et al., 1994] with 3-D interpretations. Such studies will greatly assist scientists in their understanding of the dimensions, position, and complexity of the conduit and magmatic system. Further, knowledge of the physical makeup of the volcano's edifice will help to generate a unified model of Erebus volcano and to solve questions about the mechanism, location, and generation of magma convection and seismic signals.

Experimental and Geological Context

Erebus volcano, located just off the Ross Ice Shelf on Ross Island, is a perfect, natural laboratory for seismology, with very low seismic noise, an absence of human disturbances, ease of access, and a wide range of seismic signals (including explosions, tremor, and long- and very-long-period seismic activity). Eruptive activity at Erebus has been monitored since 1980 with a short-period seismic network and since 2003 with broadband seismometers and associated instruments (e.g., infrasound and Global Positioning System (GPS) receivers; tilt-meters; and video, thermal, and meteorological sensors [see Aster et al., 2004]). Ongoing multidisciplinary studies of Erebus [Oppenheimer and Kyle, 2008] evidence the complexity of the volcano's evolution.

These studies have revealed that the Erebus stratovolcano began forming 1.3 million years ago as part of the larger Erebus volcanic province, at the south end of the Terror rift (an intraplate rift that is a major arm of the West Antarctic rift). Seismic studies suggest that the mantle beneath Erebus is 150–200 (± 100) kelvins hotter than its surroundings, showing the existence of a thermal anomaly under the region [Watson et al., 2006]. The surface expression of this thermal anomaly is a

lava lake that is phonolitic, meaning that the magma feeding Mount Erebus is alkaline rich, a factor indicative of intraplate magmatism.

Tomo Erebus Setup: Determining Seismic Array Geometry

The goals of TE are to obtain a P wave 2-D velocity profile for Ross Island and a 3-D velocity model for the Erebus summit. The 3-D velocity model will apply the tomography code of Toomey et al. [1994], already used in local active-source seismic tomography experiments at mid-ocean ridges and volcanic islands [e.g., Zandomenighi et al., 2009]. This method, based on P wave travel times, exploits a shortest-time ray tracing and a least squares algorithm inversion. An advantage of the code is that it takes into account even steep topography, allowing scientists to model seismic ray propagation and velocity in fine detail.

Prior to fieldwork, the code was run to test different code parameters and to establish the optimal geometry for shots and receivers such that scientists could best image the volcano's structure. The synthetic tests consisted of assuming different configurations of sources and receivers in a known 3-D velocity model. After calculating the travel times for each setting, the obtained data set was inverted and the final velocity structure was compared with the known model. Good performance was shown when virtual nodes within the model were spaced 50 meters apart to optimize ray tracing and 100 meters apart to fine tune velocity signatures. From this array of virtual nodes, different geometries for stations and shots were tested. On the basis of this, scientists decided to deploy 100 stations specifically meant to image the structure of Erebus in 3-D over 9 square kilometers, with an interstation distance of 350–500 meters.

Tomo Erebus Experiment

During the 2007–2008 austral summer season, 23 intermediate-period three-component seismic sensors (natural frequency 30 seconds to 100 hertz) were deployed for the passive portion of the experiment. These sensors recorded natural seismic activity during the 2008 winter. To prepare for the harsh and isolated conditions of winter, seismic stations were specially designed by the Incorporated Research Institutions for Seismology (IRIS) Program for Array Seismic Studies of the Continental Lithosphere (PASSCAL) Instrument Center for Antarctic weather conditions (<http://www.passcal.nmt.edu/Polar/index.html>). For each station this design included a 24-bit data acquisition system (DAS, with a sampling rate of 40 samples per second), a solar power controller, and batteries specifically designed for long life housed in insulated boxes. Each station also included a GPS antenna and two 65-watt solar panels mounted on reinforced aluminum frames.

The 2007–2008 fieldwork also provided an opportunity to scope out where the active seismic experiment would take place during the following summer season. Using chemical blasts as seismic sources, because natural seismic activity is low and the Antarctic summer field season is short, scientists tested three trial locations. Armed with these data, scientists in the 2008–2009 summer field season split their work into two separate experiments (Figure 1): TE-2D and TE-3D. In TE-2D, scientists deployed 21 seismometers along a 90-kilometer-long east-west line across Ross Island, using three chemical blasts to image the 2-D deep structure of the region. One shot each was positioned at both ends of the array, and a third shot was placed in the array's middle.

In TE-3D, scientists arranged 79 seismometers in a 3 × 3-kilometer grid centered on the summit crater to obtain a 3-D model of the conduit and of the upper few kilometers beneath the lava lake. One hundred short-period three-component sensors (4.5 hertz) were used to collect data for TE-2D and TE-3D, with a sampling rate of 200 samples per second. These short-period stations held one 10-watt solar panel and either a 58- or 32-ampere-hour gel-cell battery. The same IRIS setup of insulated boxes was used, although the solar panel and GPS antenna were mounted directly on top of the boxes for the TE-3D stations. The 13 shots for TE-3D were azimuthally distributed to provide optimal seismic ray path coverage (Figure 1).

The blasts themselves were made by first drilling holes 20 centimeters in diameter in snow

and ice to depths of 7–15 meters. After being loaded with explosives, the holes were stemmed with snow and left at least 24 hours to sinter. Scientists used an ammonium nitrate and fuel oil mixture (ANFO) with pentolite boosters (a mixture of pentaerythritol tetranitrate and trinitrotoluene) for blasting with detonation cord and seismic detonators. Shots ranged in size between 75 and 230 kilograms for the 2007–2008 tests and between 75 and 600 kilograms in 2008–2009. Dynamite was used for a shot in McMurdo Sound. Shot locations were determined using dual-frequency GPS receivers with differential methods.

Data Analysis to Build a Model of Erebus's Conduit

In a preliminary examination of DAS data and log files, which describe each station's state of health, GPS performance, power input, sensor channel centering, and timing errors, only very few pieces of data were found to be missing, presumably due to malfunctions of instruments. Shots for TE-3D were recorded over the entire summit array and some of the flank stations with very low noise (see Figure S1 in the electronic supplement to this Eos issue (http://www.agu.org/eos_elec/)). Shot records of TE-2D are more uncertain, but most stations on the east side of the line deployment recorded the westernmost shot with high quality. Initial data analysis indicates that waveforms and frequencies of signals recorded vary across the network, suggesting medium heterogeneity. Nonetheless, the first arrival times of P waves should be identifiable with small uncertainty, helping to better model the traveltime information.

Data recorded by the stations installed in 2007–2008 for passive seismic analysis will be introduced in this tomography to increase lateral resolution and modeled depth. In addition, it is hoped that the 2-D data will help to delineate the presence of major velocity anomalies related to the thermal imprint of what may be wide-scale mantle unconformity. Because of its higher resolution, TE-3D is expected to image the presence of hectometer-scale velocity perturbations in the first 2 kilometers of the volcano edifice, due to volcanic deposits, thermal anomalies, geochemically altered rocks, and magma bodies. Both 2-D and 3-D data will help scientists better understand the regional tectonics and its local volcanic implications.

Acknowledgments

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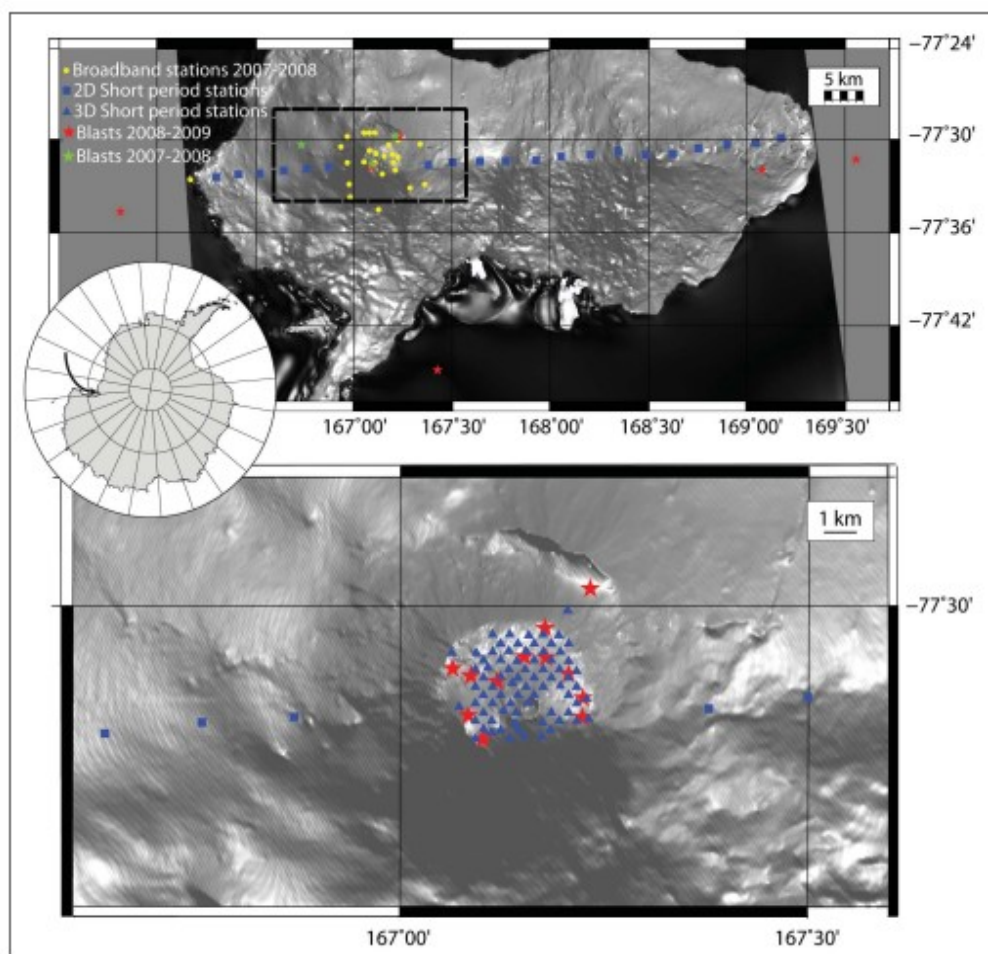


Fig. 1. Top map and inset show Ross Island and the experiment configuration for the 2007–2008 and 2008–2009 field seasons. Bottom image shows a close-up of the locations of instruments deployed on Mount Erebus's summit.

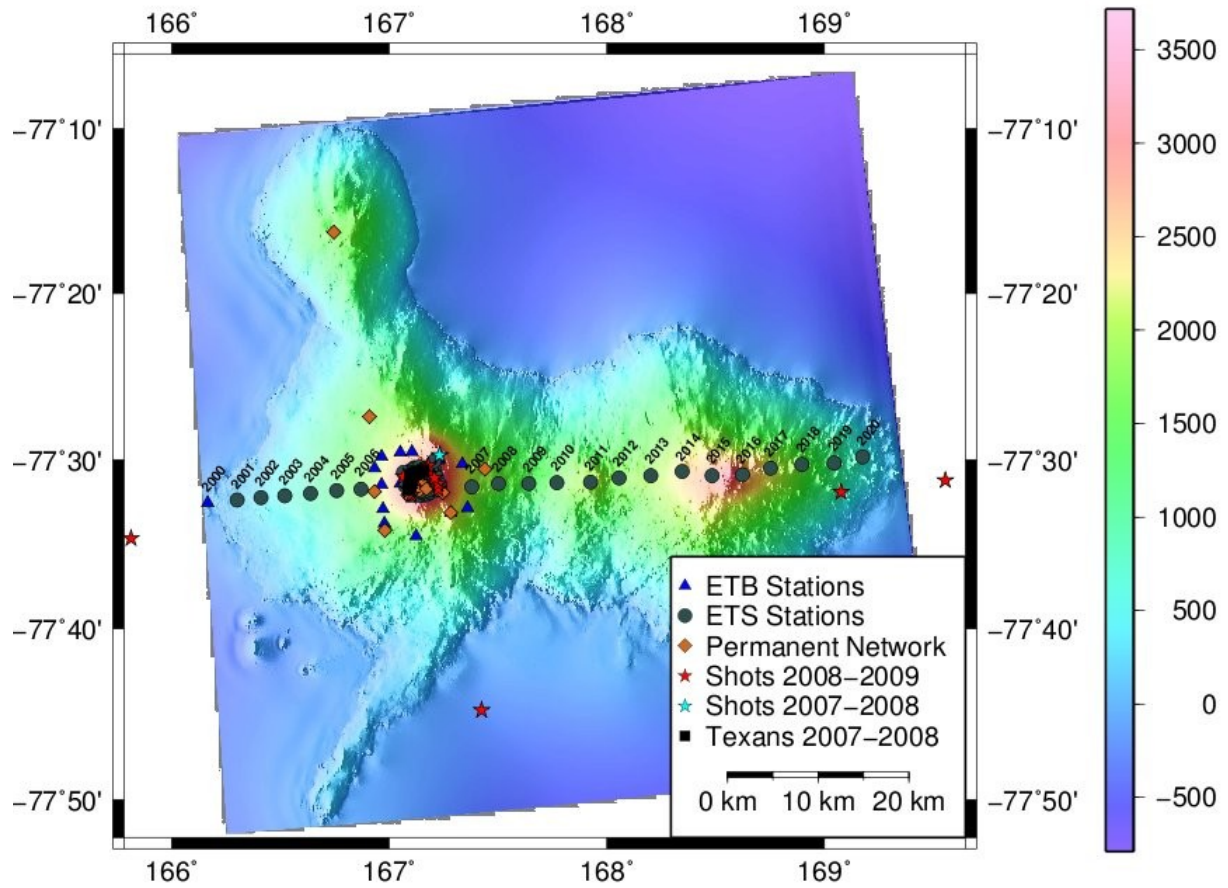


Figure 2: Map of Ross Island showing all stations and shots from the 2007-2008, and 2008-2009 field seasons.

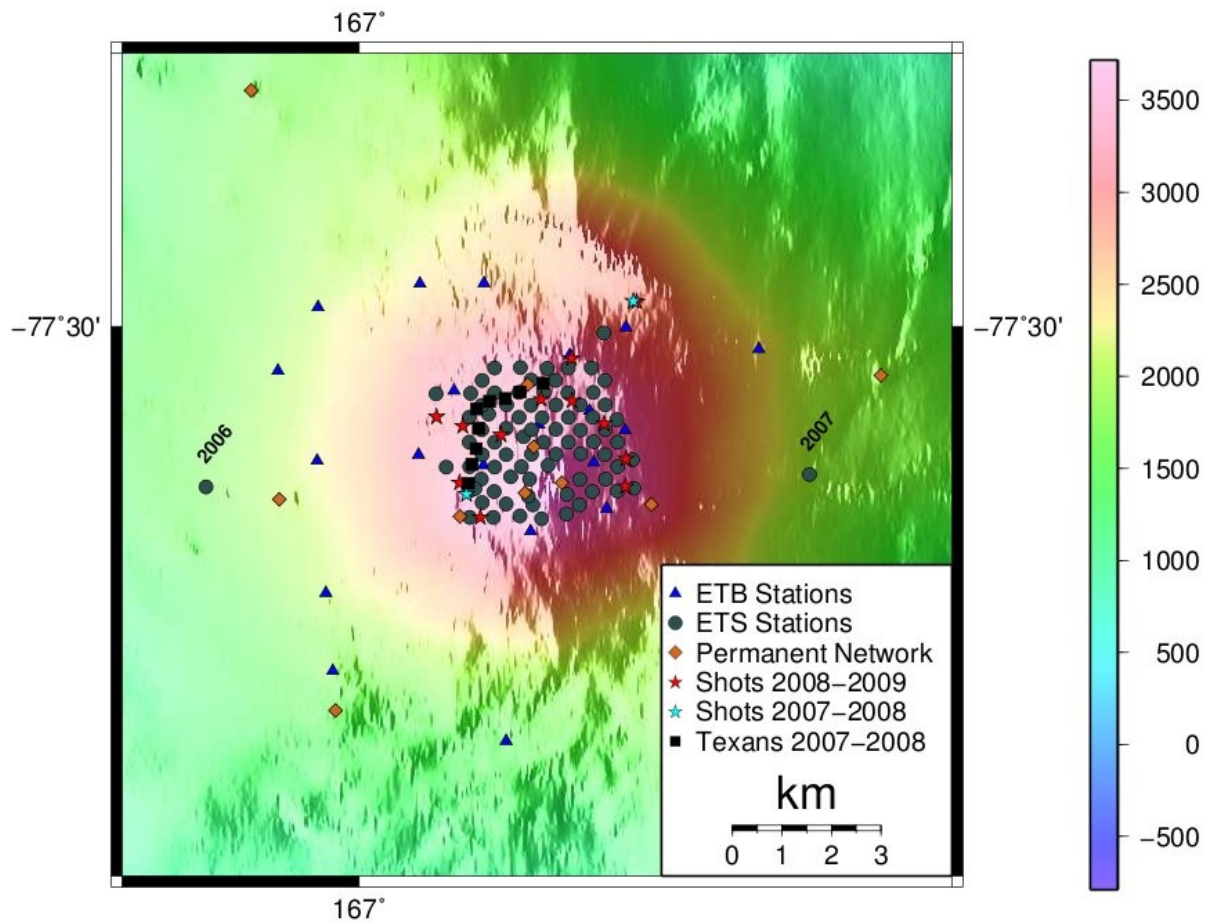


Figure 3: Close up of the summit showing all stations and shots from the 2007-2008 and 2008-2009 field seasons.

| Shot Point ID | Shot Point Name | Time (UTC) yyyy:jd:hr:mn:ss | latitude (WGS84) | longitude (WGS84) | elevation (m) | type of source | size (kg) | depth below surface (m) | # of holes |
|---------------|-----------------------------|--------------------------------|---------------------|----------------------|------------------|-------------------|-----------|----------------------------|---------------|
| 4000 | Windless Bight | 2008:346:01:40:00 | 77.74667 | 167.42534 | 38.774 | ANFO | 500 | 15 | 4 |
| 4001 | Cape Crozier | 2008:346:00:17:00 | 77.52047 | 169.55677 | 23.408 | ANFO | 600 | 14.5 | 5 |
| 4002 | Cape Royd | 2008:345:05:17:00 | 77.57793 | 165.81121 | -0.123 | dynamite | 200 | water | 1 |
| 4003 | Crozier 2 | 2008:357:02:27:00 | 77.53207 | 169.07915 | 792.162 | ANFO | 500 | 15 | 5 |
| 4004 | FANG | 2008:355:21:51:00 | 77.49550 | 167.23412 | 2928.494 | ANFO | 300 | 15 | 3 |
| 4005 | Cones2 | 2008:356:09:00:00 | 77.53486 | 167.10207 | 3494.562 | ANFO | 100 | 6.0, 6.13, 7.4 | 3 |
| 4006 | Cones | 2008:356:08:23:00 | 77.52857 | 167.08469 | 3439.634 | ANFO | 200 | 14.9 | 2 |
| 4007 | CORR13 | 2008:356:07:50:00 | 77.51817 | 167.08751 | 3294.538 | ANFO | 75 | 20 | 1 |
| 4008 | Tramsw 2 | 2008:356:09:45:00 | 77.51977 | 167.11968 | 3421.712 | ANFO | 75 | 6.15, 6.8 | 2 |
| 4009 | Sunshine Valley CornerSW | 2008:359:00:30:00 | 77.51650 | 167.06509 | 3225.665 | ANFO | 100 | 5.5, 5.8, 7.9 | 3 |
| 4010 | HoleH | 2008:358:20:55:00 | 77.51334 | 167.15366 | 3424.817 | ANFO | 75 | 7.8, 8.15 | 2 |
| 4011 | Stinky (13) | 2008:358:23:49:00 | 77.51361 | 167.17945 | 3424.82 | ANFO | 75 | 8.1, 8.5 | 2 |
| 4012 | Black (19) | 2008:357:05:02:00 | 77.52920 | 167.22458 | 3461.882 | ANFO | 100 | 7.5 | 2 |
| 4013 | Tower (17) | 2008:357:05:28:00 | 77.52416 | 167.22520 | 2928.494 | ANFO | 100 | 7.8, 8.7 | 2 |
| 4014 | Fog (15) | 2008:357:05:55:00 | 77.51781 | 167.20687 | 3466.01 | ANFO | 100 | 8.0, 8.6 | 2 |
| 4015 | Stuck (11) | 2008:358:23:27:00 | 77.50585 | 167.17907 | 3351.054 | ANFO | 100 | 6.2 7.1 | 2 |

| Receiver ID | latitude (WGS84) | longitude (WGS84) | Elevation (m) | Serial # of data logger | data logger manufacturer | data logger model # | sensor manufacturer | sensor model number |
|-------------|---------------------|----------------------|------------------|-------------------------------|-----------------------------|------------------------|------------------------|---------------------------|
| 1001 | -77.5245 | 166.964417 | 2359 | 9297 | RefTek | RT-130 | Guralp | 40-T |
| 1002 | -77.548567 | 166.97205 | 2114 | 9892 | RefTek | RT-130 | Guralp | 40-T |
| 1003 | -77.508133 | 166.931617 | 2001 | 9866 | RefTek | RT-130 | Guralp | 40-T |
| 1004 | -77.4965 | 166.965167 | 2143 | 9873 | RefTek | RT-130 | Guralp | 40-T |
| 1005 | -77.492167 | 167.051167 | 2452 | 9848 | RefTek | RT-130 | Guralp | 40-T |
| 1006 | -77.492083 | 167.105167 | 2583 | 9915 | RefTek | RT-130 | Guralp | 40-T |
| 1007 | -77.5628 | 166.9777 | 1780 | 995F | RefTek | RT-130 | Guralp | 40-T |
| 1008 | -77.504183 | 167.336983 | 2495 | 990B | RefTek | RT-130 | Guralp | 40-T |
| 1009* | -77.542717 | 166.1646 | 16 | 995D | RefTek | RT-130 | Guralp | 40-T |
| 1010 | -77.55235 | 167.282717 | 2361 | 92C8 | RefTek | RT-130 | Guralp | 40-T |
| 1011 | -77.517917 | 167.151567 | 3494 | 976C | RefTek | RT-130 | Guralp | 40-T |
| 1012 | -77.515117 | 167.109217 | 3373 | 985B | RefTek | RT-130 | Guralp | 40-T |
| 1013 | -77.547933 | 167.360350 | 1979 | 9868 | RefTek | RT-130 | Guralp | 40-T |
| 1014 | -77.515417 | 167.194300 | 3437 | 944B | RefTek | RT-130 | Guralp | 40-T |
| 1015 | -77.537333 | 167.144517 | 3405 | 9859 | RefTek | RT-130 | Guralp | 40-T |
| 1016 | -77.511700 | 167.079967 | 3274 | 92D9 | RefTek | RT-130 | Guralp | 40-T |
| 1017 | -77.533283 | 167.208633 | 3437 | 995A | RefTek | RT-130 | Guralp | 40-T |
| 1018 | -77.524883 | 167.197683 | 3566 | 995B | RefTek | RT-130 | Guralp | 40-T |
| 1019 | -77.505250 | 167.177533 | 3290 | 988F | RefTek | RT-130 | Guralp | 40-T |
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| 1021 | -77.500167 | 167.225217 | 2951 | 984D | RefTek | RT-130 | Guralp | 40-T |
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| 1023 | -77.523383 | 167.050150 | 3236 | 9343 | RefTek | RT-130 | Guralp | 40-T |
| 1024 | -77.575517 | 167.124017 | 1540 | 9876 | RefTek | RT-130 | Guralp | 40-T |

* Sensor changed to Mark Products L-28 on November 10th 2008.

| Receiver ID | latitude (WGS84) | longitude (WGS84) | Elevation (m) | serial # of data logger | data logger manufacturer | data logger model # | sensor manufacturer | sensor model number |
|-------------|------------------|-------------------|---------------|-------------------------|--------------------------|---------------------|---------------------|---------------------|
| 2001 | -77.53994 | 166.30038 | 173 | 9777 | RefTek | RT-130 | Mark Products | L-28 |
| 2002 | -77.53737 | 166.40999 | 398 | 9142 | RefTek | RT-130 | Mark Products | L-28 |
| 2003 | -77.5357 | 166.52064 | 633 | 9095 | RefTek | RT-130 | Mark Products | L-28 |
| 2004 | -77.53352 | 166.63891 | 921 | 9914 | RefTek | RT-130 | Mark Products | L-28 |
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| 2006 | -77.52926 | 166.87078 | 1680 | 9553 | RefTek | RT-130 | Mark Products | L-28 |
| 2007 | -77.52702 | 167.37993 | 2091 | 9310 | RefTek | RT-130 | Mark Products | L-28 |
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| 2015 | -77.51547 | 168.48471 | 2925 | 92BE | RefTek | RT-130 | Mark Products | L-28 |
| 2016 | -77.51452 | 168.62449 | 2860 | 924C | RefTek | RT-130 | Mark Products | L-28 |
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| 2018 | -77.50433 | 168.89784 | 1841 | 9342 | RefTek | RT-130 | Mark Products | L-28 |
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| 3079 | -77.5324 | 167.12409 | 3557 | 9334 | RefTek | RT-130 | Mark Products | L-28 |
| 3080 | -77.5346 | 167.13564 | 3521 | 9791 | RefTek | RT-130 | Mark Products | L-28 |
| 3081 | -77.53256 | 167.14668 | 3626 | 980E | RefTek | RT-130 | Mark Products | L-28 |
| 3082 | -77.50983 | 167.12429 | 3529 | 986C | RefTek | RT-130 | Mark Products | L-28 |
| 3083 | -77.51224 | 167.09409 | 3331 | 9C30 | RefTek | RT-130 | Mark Products | L-28 |
| 3084 | -77.51233 | 167.06478 | 3235 | 943F | RefTek | RT-130 | Mark Products | L-28 |
| 3085 | -77.52435 | 167.23053 | 3458 | 9924 | RefTek | RT-130 | Mark Products | L-28 |
| 3086 | -77.52793 | 167.18616 | 3601 | 9462 | RefTek | RT-130 | Mark Products | L-28 |
| 3087 | -77.53258 | 167.1859 | 3515 | 9237 | RefTek | RT-130 | Mark Products | L-28 |
| 3088 | -77.53063 | 167.17526 | 3627 | 9926 | RefTek | RT-130 | Mark Products | L-28 |
| 3089 | -77.50752 | 167.13603 | 3369 | 9245 | RefTek | RT-130 | Mark Products | L-28 |
| 3090 | -77.53096 | 167.14265 | 3678 | 9009 | RefTek | RT-130 | Mark Products | L-28 |
| 3091 | -77.52572 | 167.07326 | 3426 | 956F | RefTek | RT-130 | Mark Products | L-28 |
| 3092 | -77.52332 | 167.14515 | 3658 | 925D | RefTek | RT-130 | Mark Products | L-28 |
| 3093 | -77.52562 | 167.21814 | 3494 | 92F4 | RefTek | RT-130 | Mark Products | L-28 |
| 3094 | -77.52788 | 167.20786 | 3515 | 909A | RefTek | RT-130 | Mark Products | L-28 |
| 3095 | -77.53015 | 167.19658 | 3527 | 9560 | RefTek | RT-130 | Mark Products | L-28 |
| 3096 | -77.53007 | 167.2178 | 3488 | 92C4 | RefTek | RT-130 | Mark Products | L-28 |
| 3097 | -77.52956 | 167.23168 | 3439 | 92AC | RefTek | RT-130 | Mark Products | L-28 |
| 3098 | -77.52537 | 167.15936 | 3654 | 991C | RefTek | RT-130 | Mark Products | L-28 |
| 3099 | -77.50123 | 167.20651 | 3037 | 92AB | RefTek | RT-130 | Mark Products | L-28 |